

Optics

- Introduction
- Layout
- Field
- Performance
 - Spot Diagram
 - Aberration
 - Ghost
 - Thermal
- Tolerance
- CCD

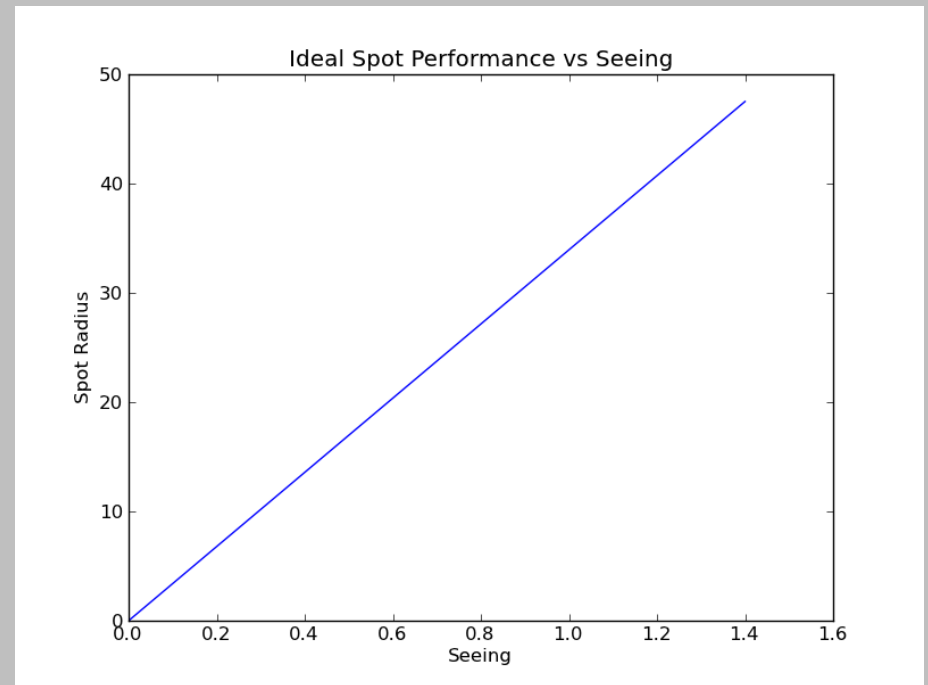
Introduction

- Focal reduction from telescope f/10.3 to f/8.0
- Spot size and ideal pixel size dictated by focal ratio
 - With a set pixel size we need to know the maximum spot size
 - Want to be Nyquist sampled, $f_s \geq 2f_c$, at best seeing, chosen to be 0.6"
 - Maximum spot radius should be 20.3 μ m

$$spot_radius = \frac{1000 * f / * aperture * f_s}{206265} \frac{\mu m}{pixel}$$

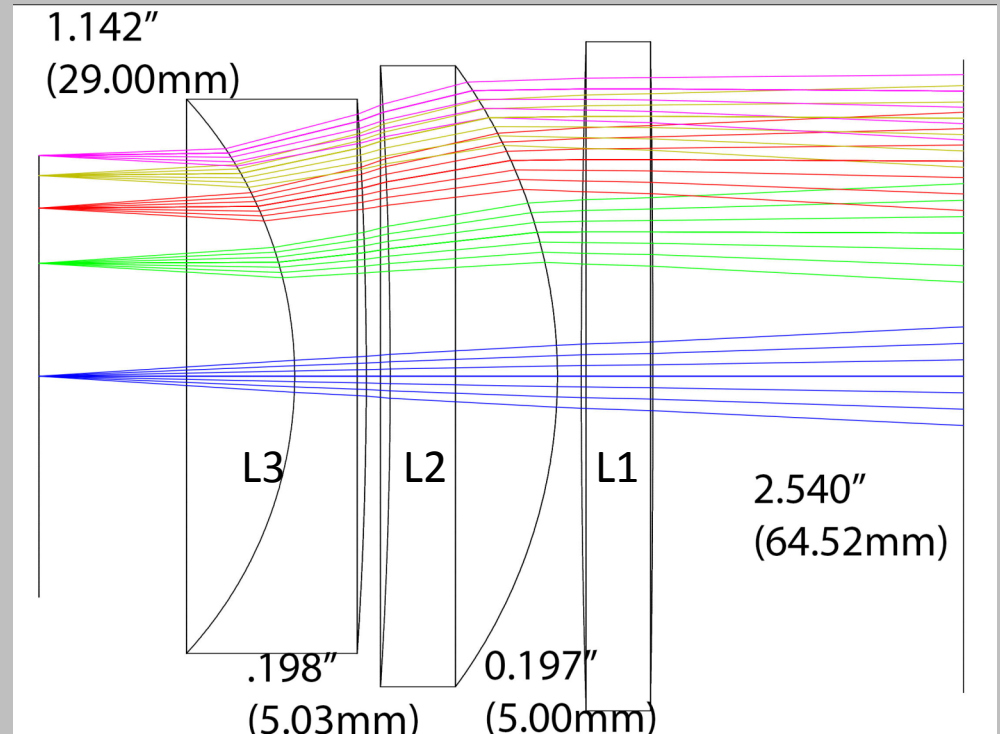
- Plate Scale will be 0.111"/pixel

$$plate_scale = \frac{206265 * 15.0 \mu m}{1000 * f / * aperture(mm)} \frac{''}{pixel}$$



Layout

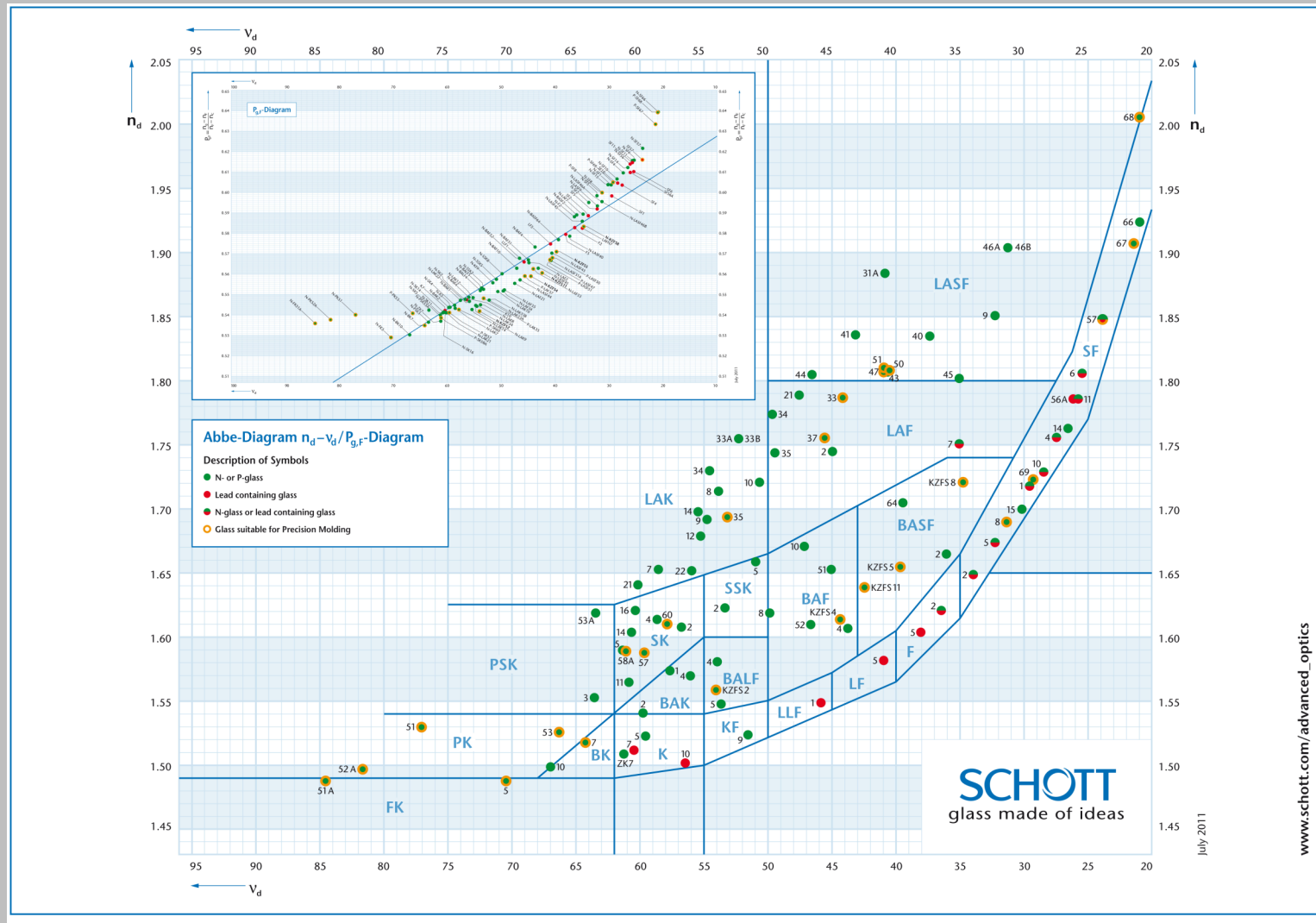
- Three element focal reducer
- Diameters with step added
 - L1 = 5.75"
 - L2 = 5.5"
 - L3 = 5.125"
- Glass types
 - L1 = SK16 (heavy crown)
 - L2 = SK14 (heavy crown)
 - L3 = F2 (flint glass)



Note: dimensions shown are minimum distance between optical surfaces (Z distance)

Crown Glass – low dispersion, low refractive index
Flint Glass – high dispersion, high refractive index

Glass Properties

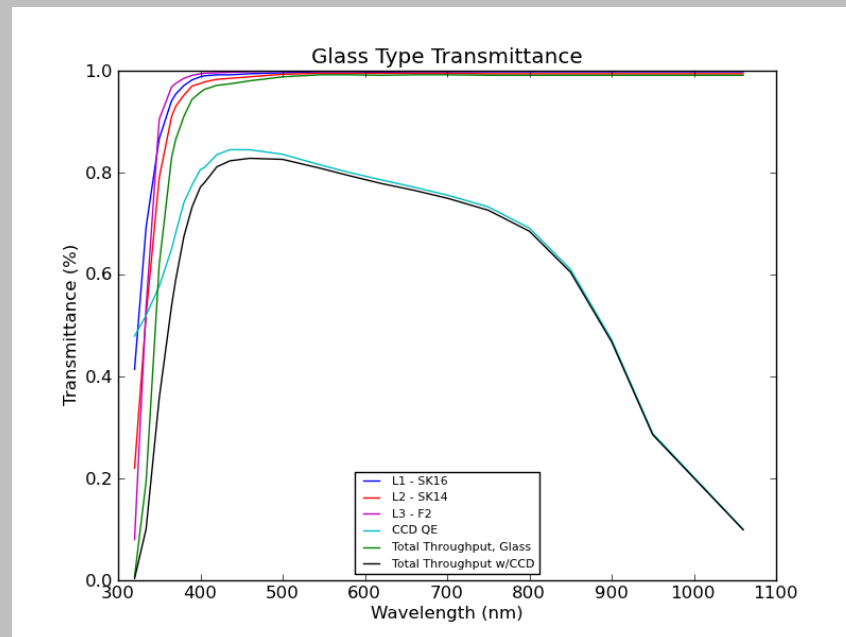


$$V_d = \frac{n_d - 1}{n_f - n_c}$$

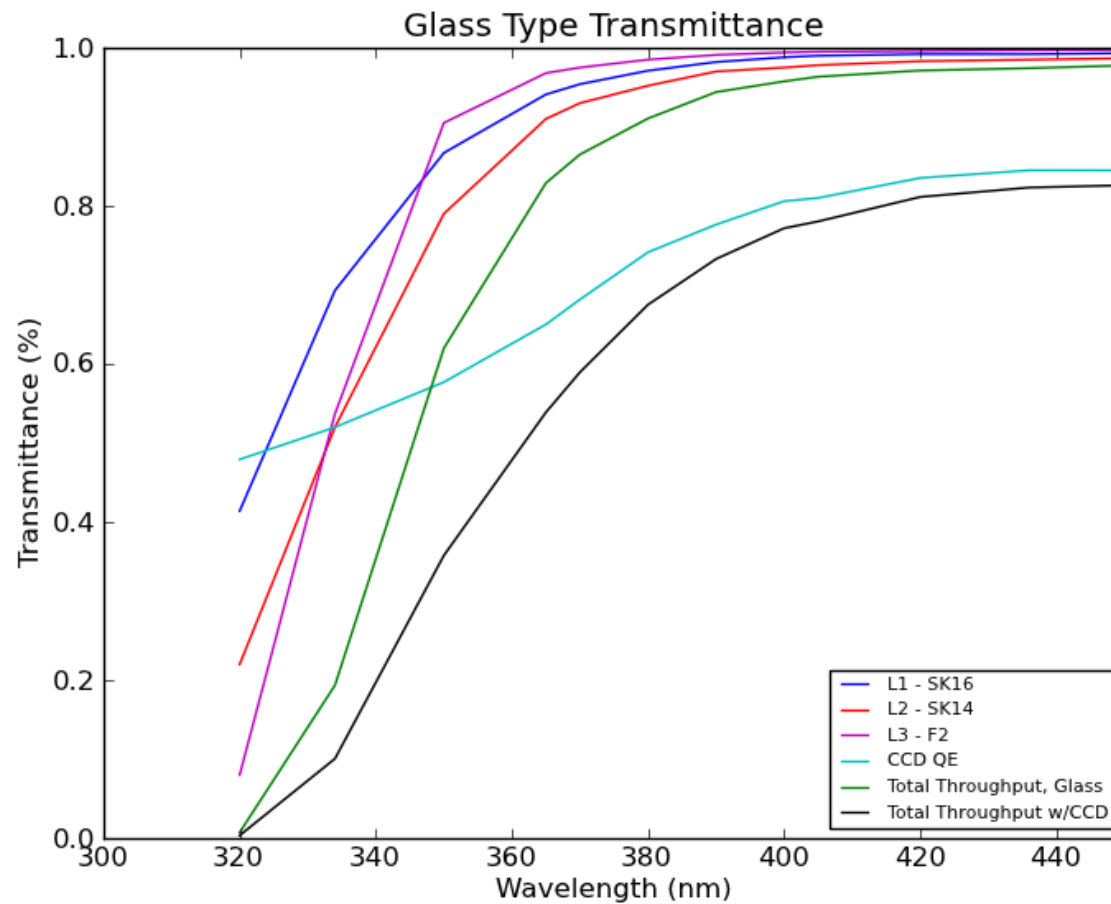
- V_d = Abbe Number is indicator of dispersion
- n_d = refractive index

Glass Selection

- Glass selection was performed by
 - creating a sub-catalog in Zemax with the selection criteria of good blue throughput and relative cost less than $2.5 * N-BK7$
 - Hammer optimization performed to reduce merit function using glass substitution along with surface variables



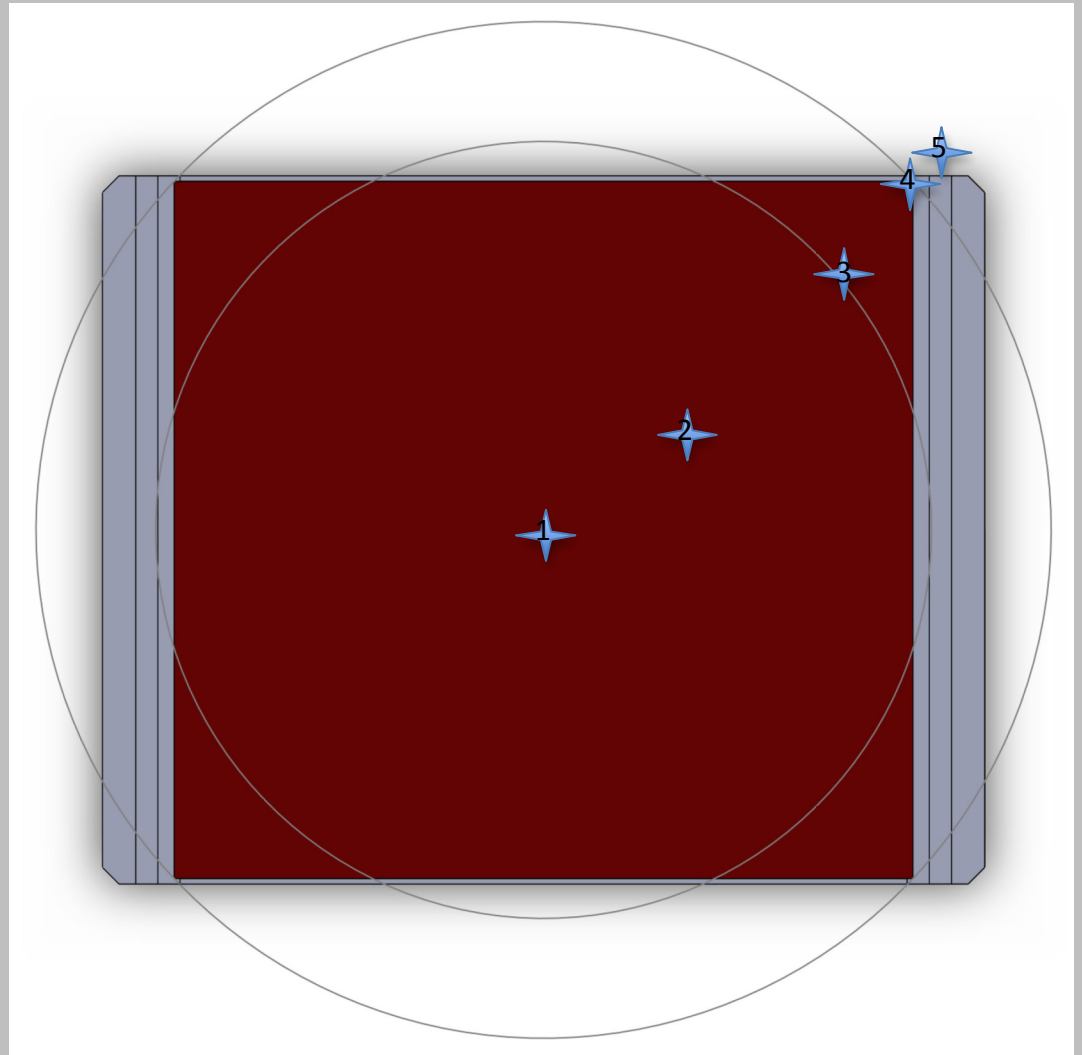
Glass Selection – Blue Performance



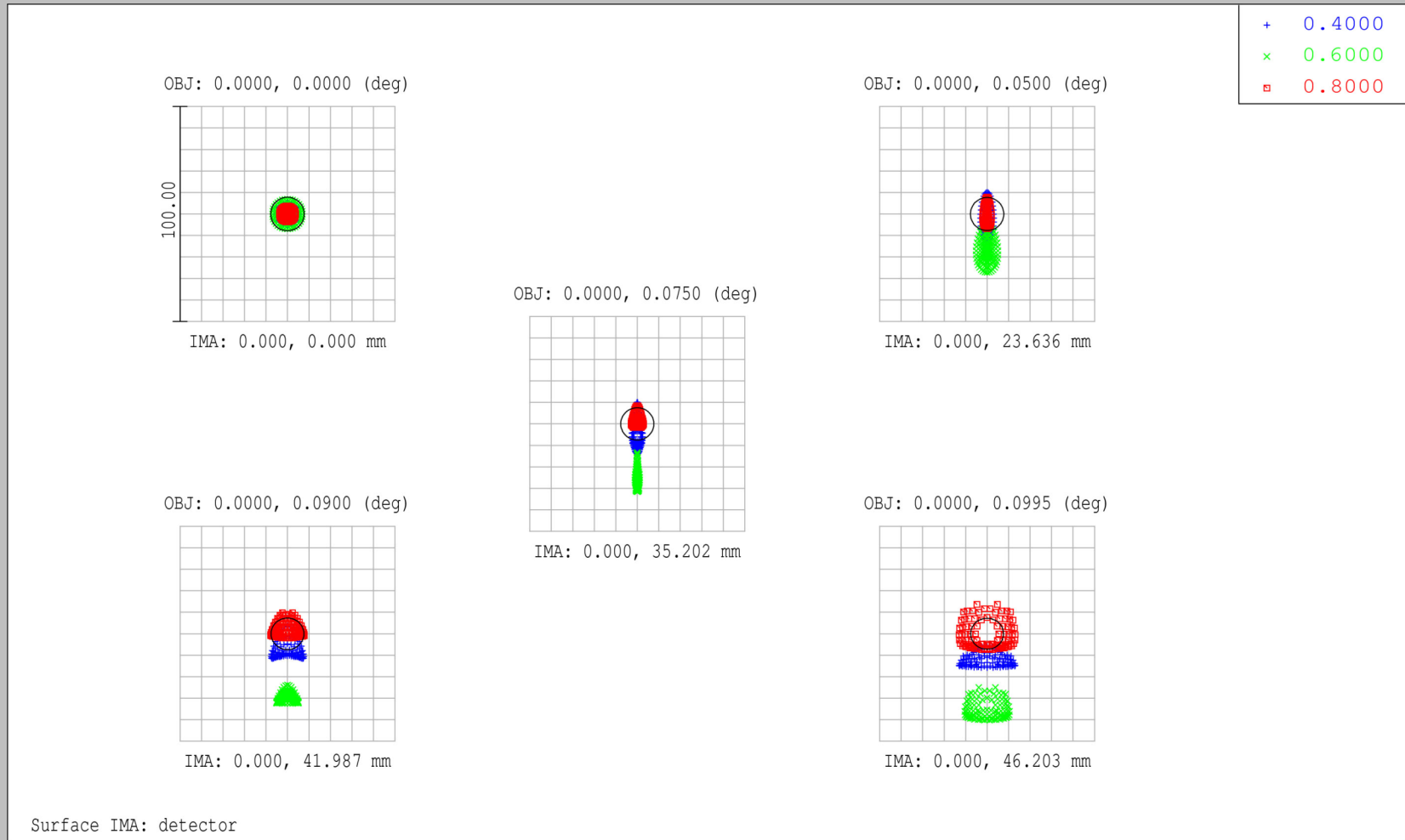
- Good response without using costly glass
- Response curves for internal transmission

Field

- Field positions sampled, in Zemax, overlayed on CCD
 - 1 = 0 deg
 - 2 = 0.05 deg (3')
 - 3 = 0.075 deg (4.5')
 - 4 = 0.09 deg (5.4')
 - 5 = 0.0995 deg (5.97')
- Radial maximum fields
 - Top edge of CCD = 3.79'
 - Right Edge of CCD = 3.80'
 - Diagonal Edge = 5.37'



Optical Performance



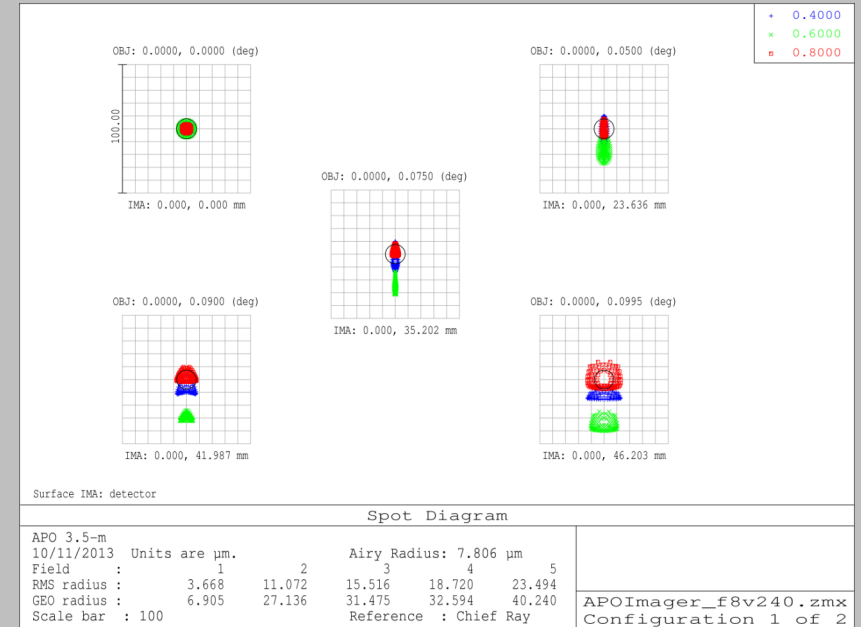
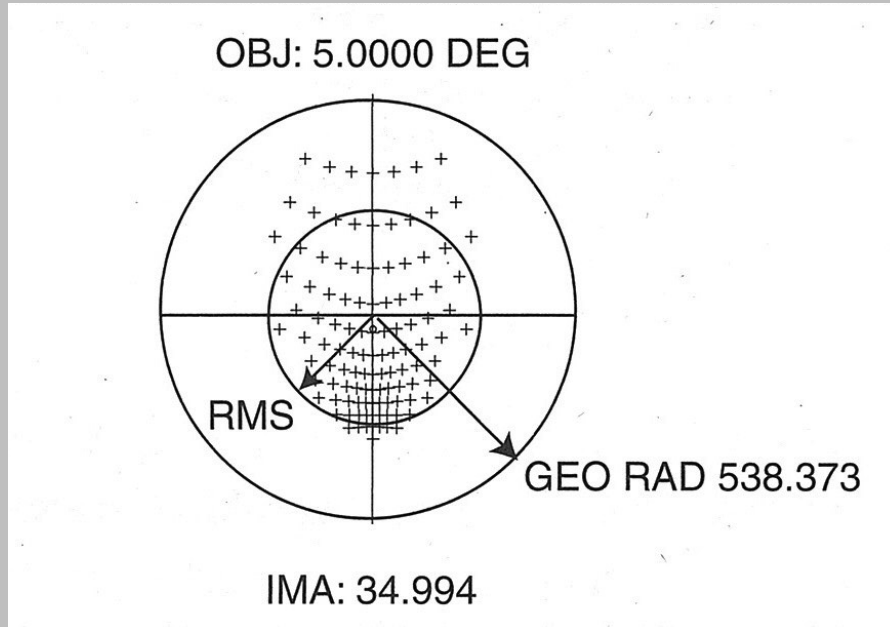
Spot Diagram

APO 3.5-m
10/11/2013 Units are μm .
Field : 1 2
RMS radius : 3.668 11.072
GEO radius : 6.905 27.136
Scale bar : 100

Airy Radius: 7.806 μm
3 4 5
15.516 18.720 23.494
31.475 32.594 40.240
Reference : Chief Ray

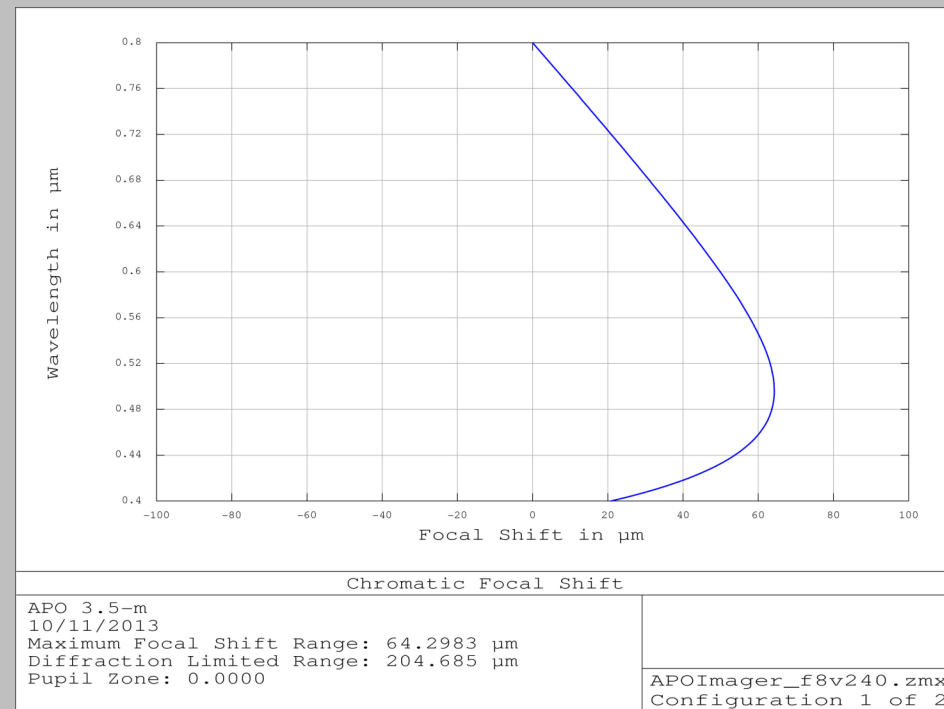
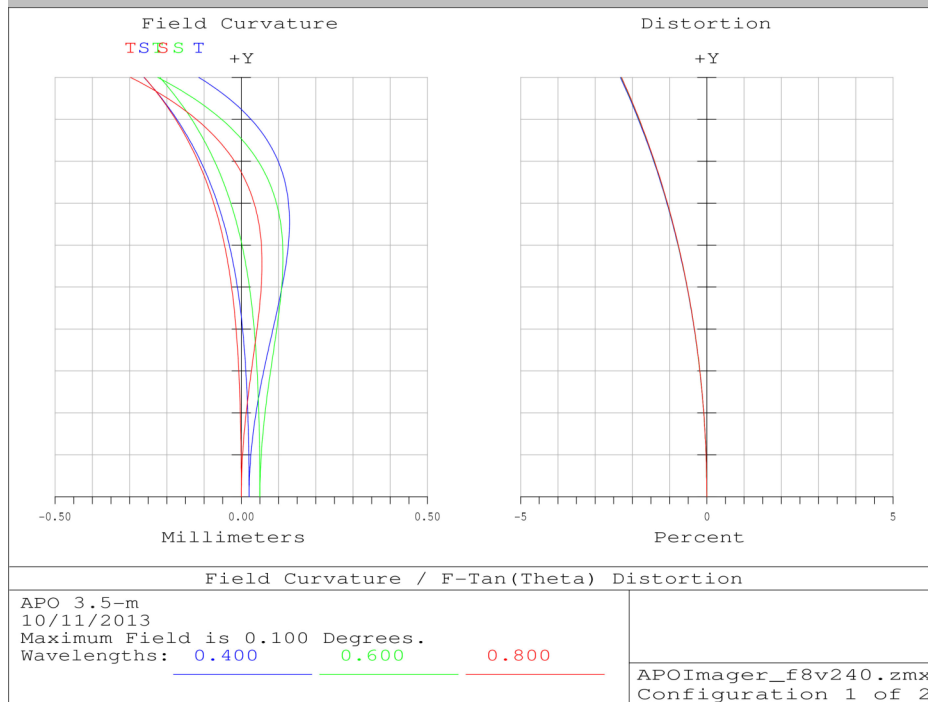
APOImager_f8v240.zmx
Configuration 1 of 2

Optical Performance



- What does the spot size mean?
- The RMS radius is the most functional number for most observers
- Geometric radius is very important for astrometric calculations
- How do these numbers relate to the real world
 - If seeing is 0.6" then to be properly sampled the spot radius needs to be 20 μm
 - If seeing is 1.0" then the spot radius needs to be 30 μm

Performance

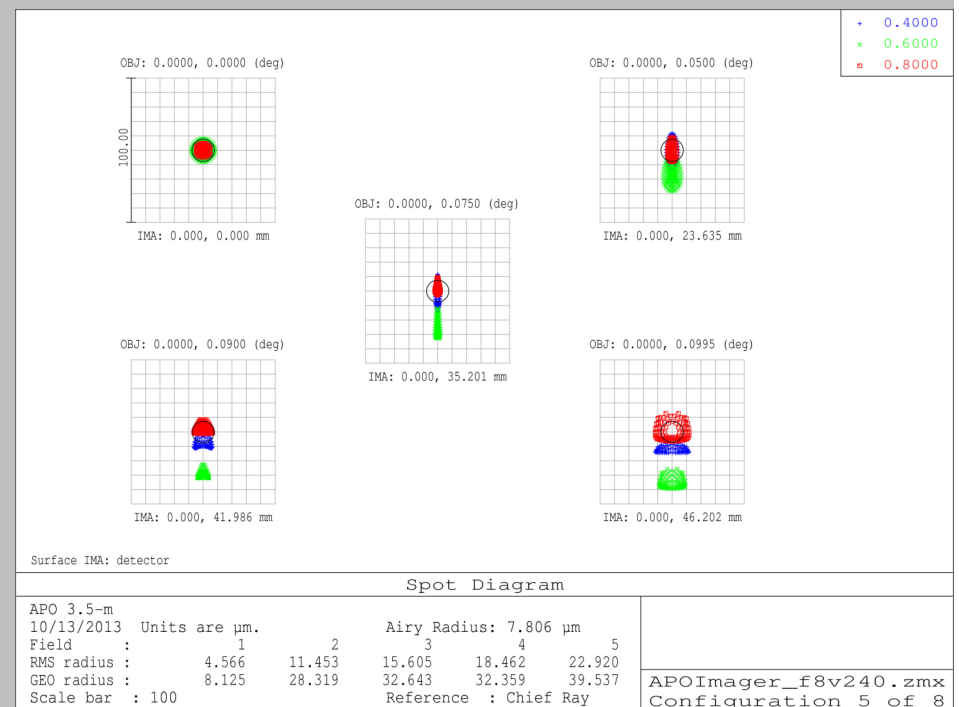
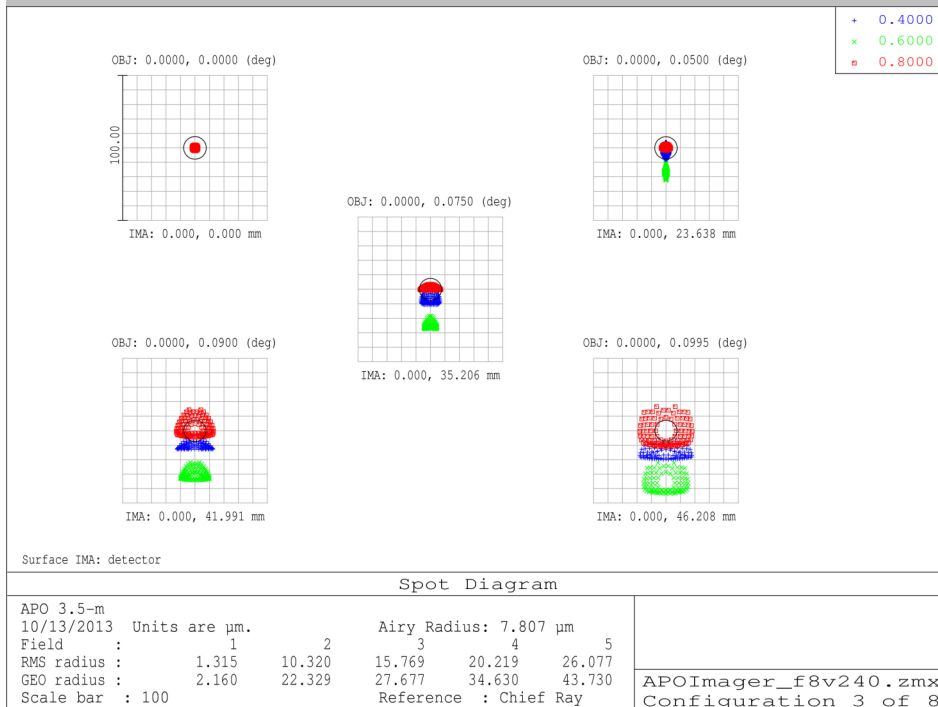


Ghost

- Ghosting optimization performed after good optical performance reached
- Ghost Pupil
 - Nearest ghost pupil located 11.1mm from CCD
 - Created off the back surface of L3 and the CCD
- Ghost Focus
 - Nearest ghost focus 13.7mm from CCD
 - Created off the back surface of L2 and the front surface of L3
- Optimal ghosting should be pushed to 18mm
- Need to perform Non-Sequential Mode Analysis to determine ghosting intensity

Thermal Modeling

- Temperature Ranges from -20C to +30C
- Temperature swings of up to 5C/hr
- Temperature changes modeled in Zemax, with spacers
- Left plot = -20C, Right plot = +30C
- Improvements seen towards center of CCD in the cold
- Improvements seen towards out edges of the CCD in warm
- Minor degradation in seeing on outer edge

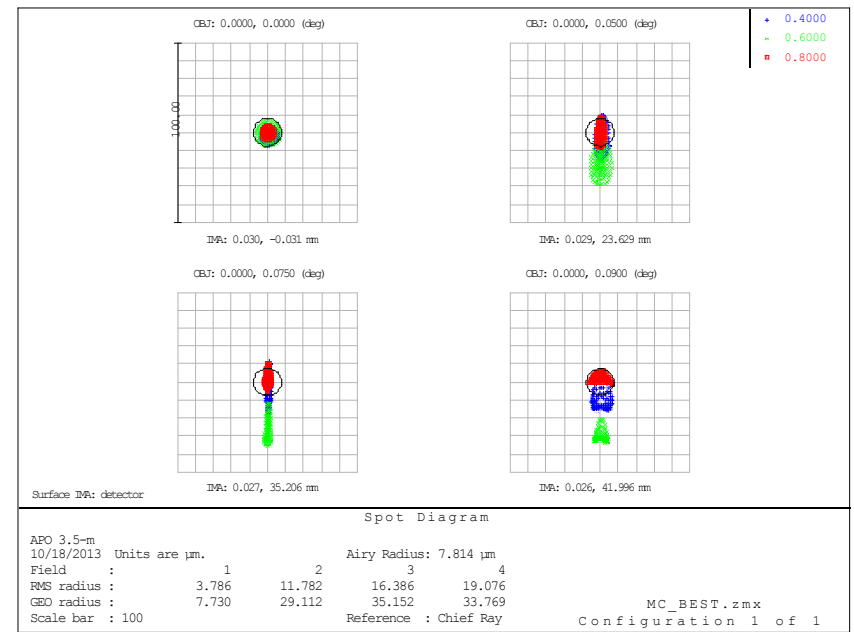
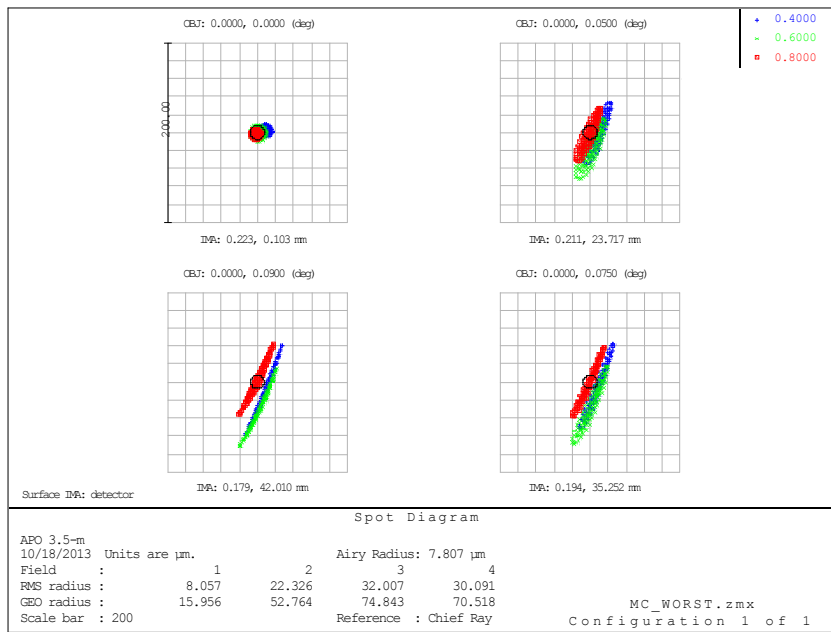


Tolerance

- Tolerancing introduces perturbations into the optical system to determine the degradation of the performance
- Surface tolerances initially set to $\pm 0.2\text{mm}$ for radius, thickness, tilt, and surface irregularities
- Element tolerance set to $\pm 0.2\text{mm}$
- Inverse Sensitivity
 - What are the parameter ranges given the maximum allowed change in performance
 - Tilt of L2 and L3 found to be the worst ‘offenders’, but still found to be acceptable
- Sensitivity
 - What is the change in performance per set tolerance

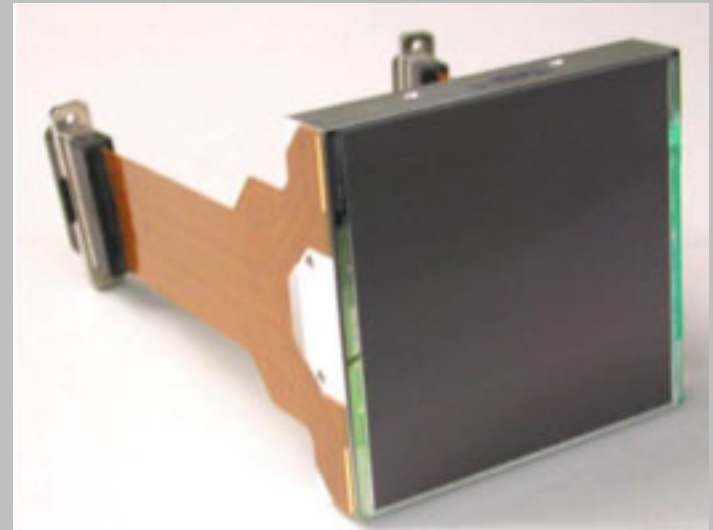
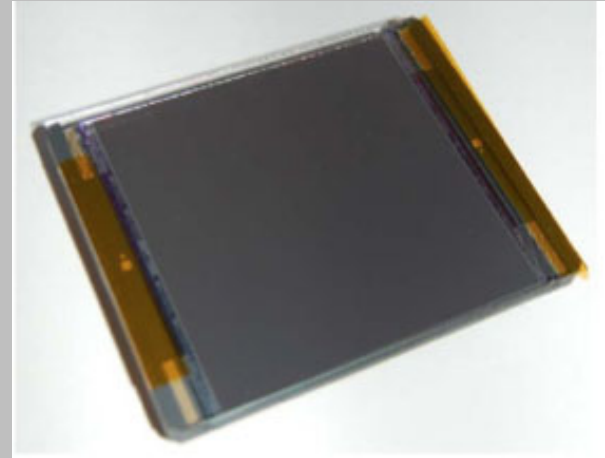
Simulation Results

- Monte Carlo simulations run on tolerancing parameters
- 41 tolerancing variables
- 100,000 simulations run



CCD

- E2V CCD 231-84
- 4K x 4K
- 15 micron pixels
- 61.4mm² imaging area
- Deep Depletion
- Ceramic Packaging



CCD Performance Comparison

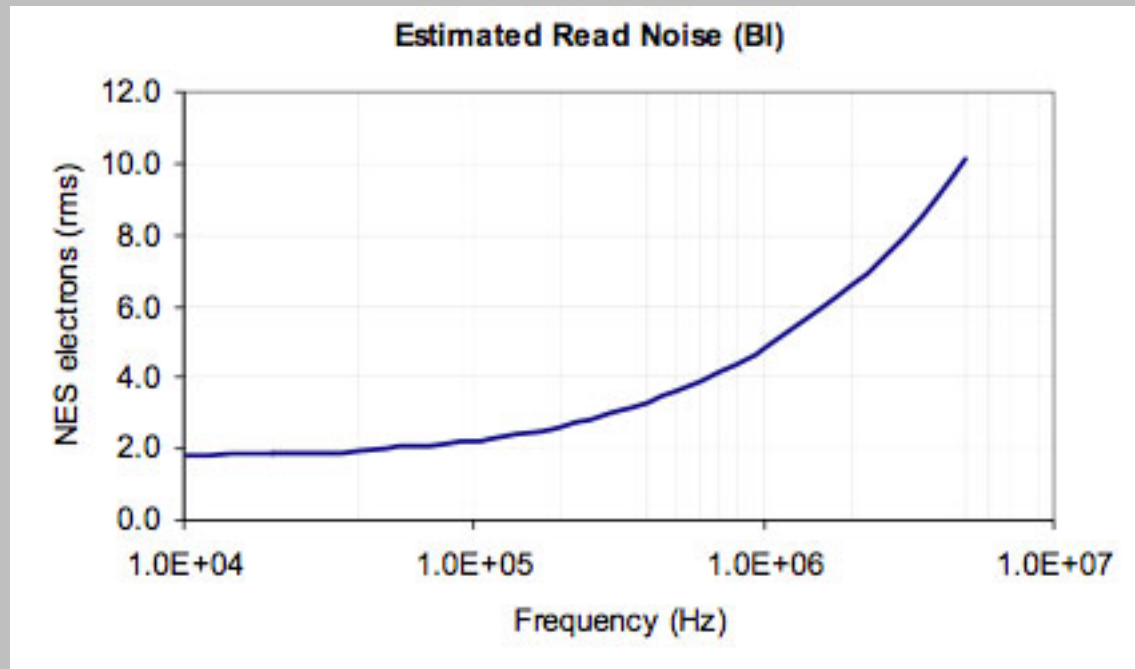
Table 2:: CCD Comparison

Spec	SPICam	CCD231
Pixel Size	24 μ m	15 μ m
Imaging Area	49.1mm ²	61.4mm ²
Read Noise	5.7e ⁻	2e ⁻ - 5e ⁻
Gain	3.36e ⁻ /ADU	2.6e ⁻ /ADU
Dark Current	2.5e ⁻ /pixel/hour	3e ⁻ /pixel/hour
Readout Rate	35kHz	1MHz - 50kHz
Readout Time	120 seconds	4.2 - 84 seconds

Notes. Readout Rate and Time only based on 1x1 binning.

- The above table shows the differences between the selected E2V CCD231-84 and SPICam
- Improvements are seen across the performance spectrum

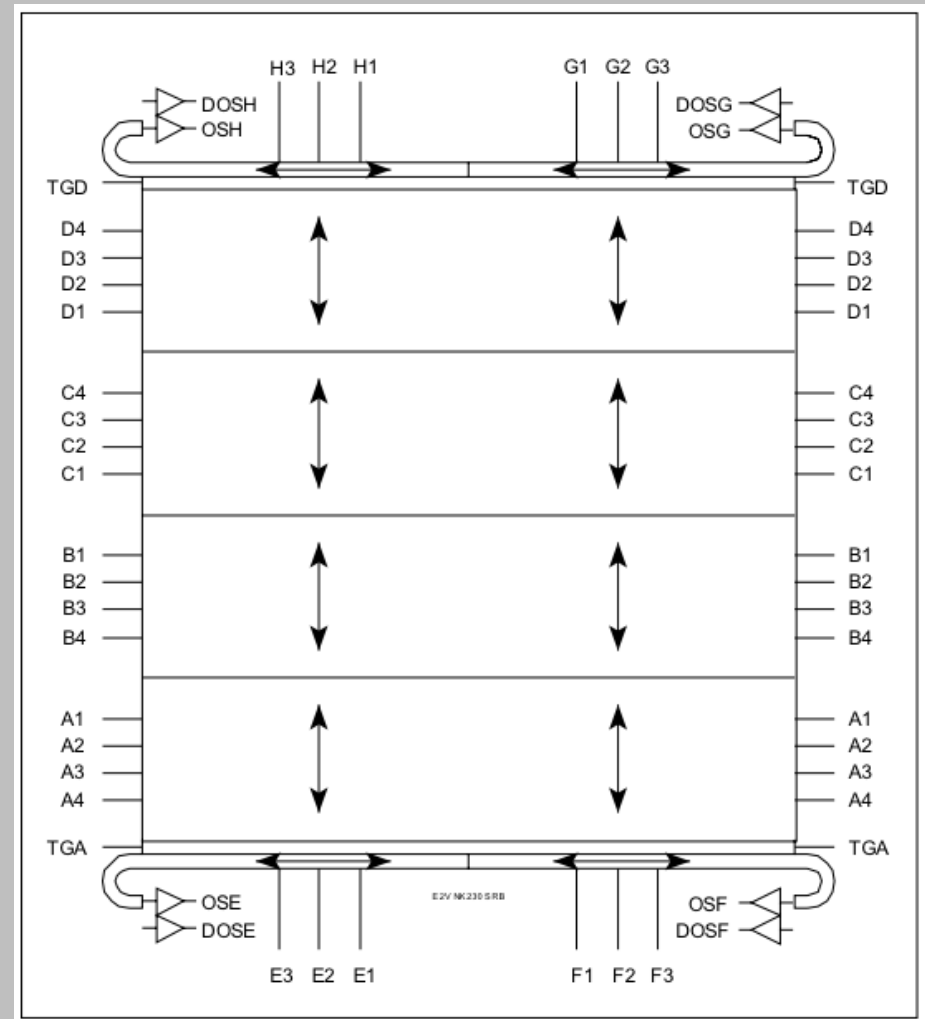
CCD Architecture



- Quad amplifier Readout
- Three read modes available (user selectable)
 - 900 KHz
 - 5 seconds readout (1x1 binning)
 - 5 electrons read noise
 - 400 KHz
 - 10 seconds (1x1 binning)
 - 3 electrons read noise
 - 100 KHz
 - 40 seconds (1x1 binning)
 - 2 electrons read noise

Readout Types

- Typical mode will use all 4 amplifiers to read out the entire chip
- User selectable command available to readout Upper two or Lower two sections independently
 - C & D clocked together
 - A & B clocked together
 - While clocking each half the other remain static
 - Possible uses include on chip guider and mosaic'ed filters

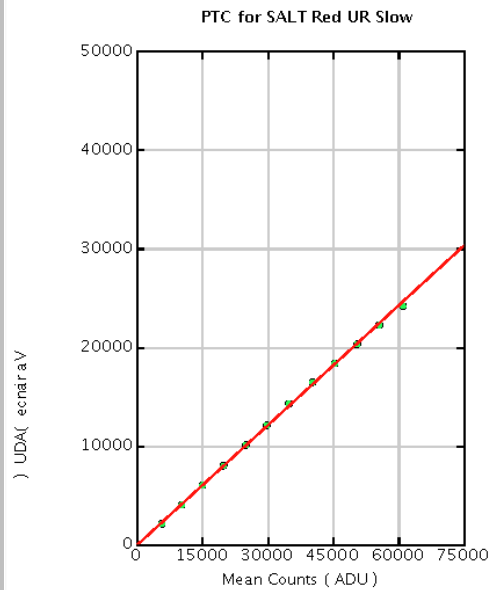


CCD Real World Examples

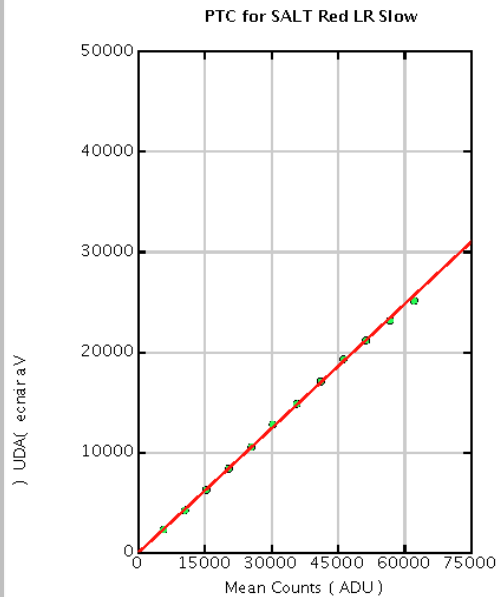
Plots are lab results of a CCD231-84 detector using Leach electronics

Left: 50KHz readout
Right: 1MHz readout

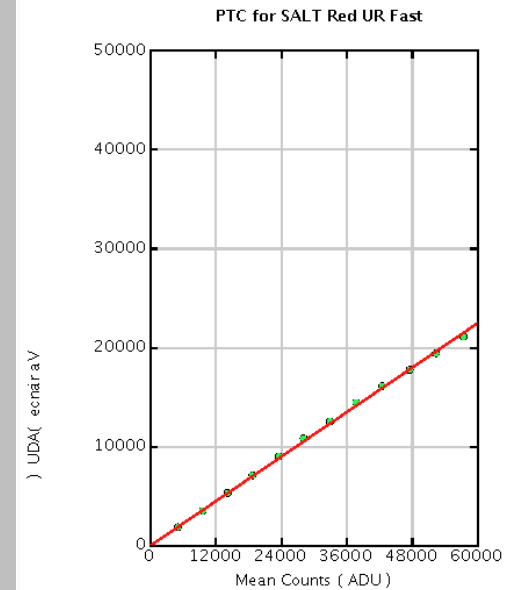
Noise results close to E2V expected numbers



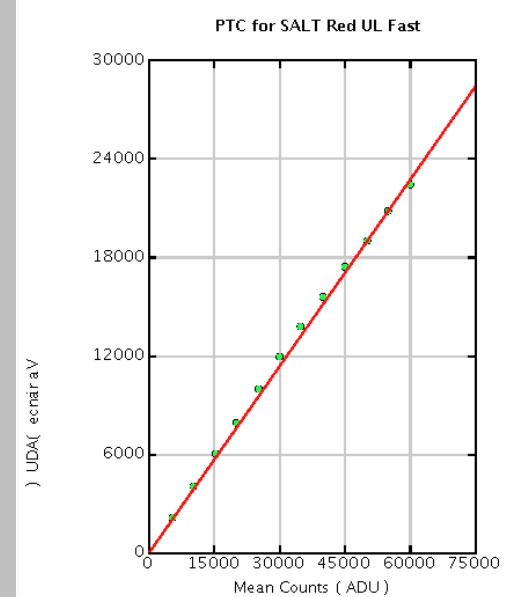
Oct 13, 2011 | 02:57 PM
Gain: 2.4672 e-/ADU
Std Dev: 1.1957 ADU
Noise: 2.95 e-



Oct 13, 2011 | 02:52 PM
Gain: 2.4167 e-/ADU
Std Dev: 1.1687 ADU
Noise: 2.8244 e-

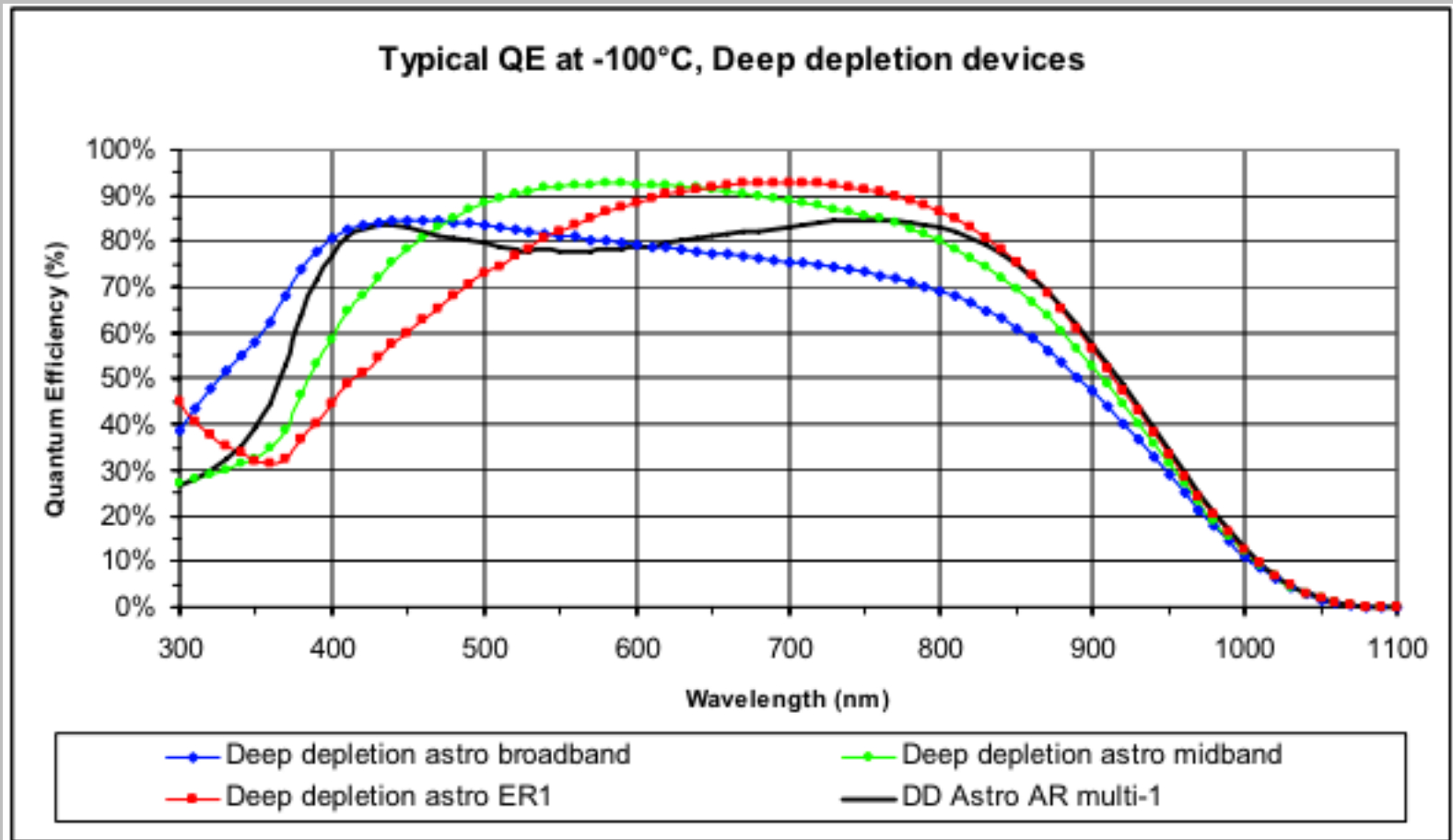


Oct 13, 2011 | 02:20 PM
Gain: 2.6663 e-/ADU
Std Dev: 1.772 ADU
Noise: 4.7247 e-



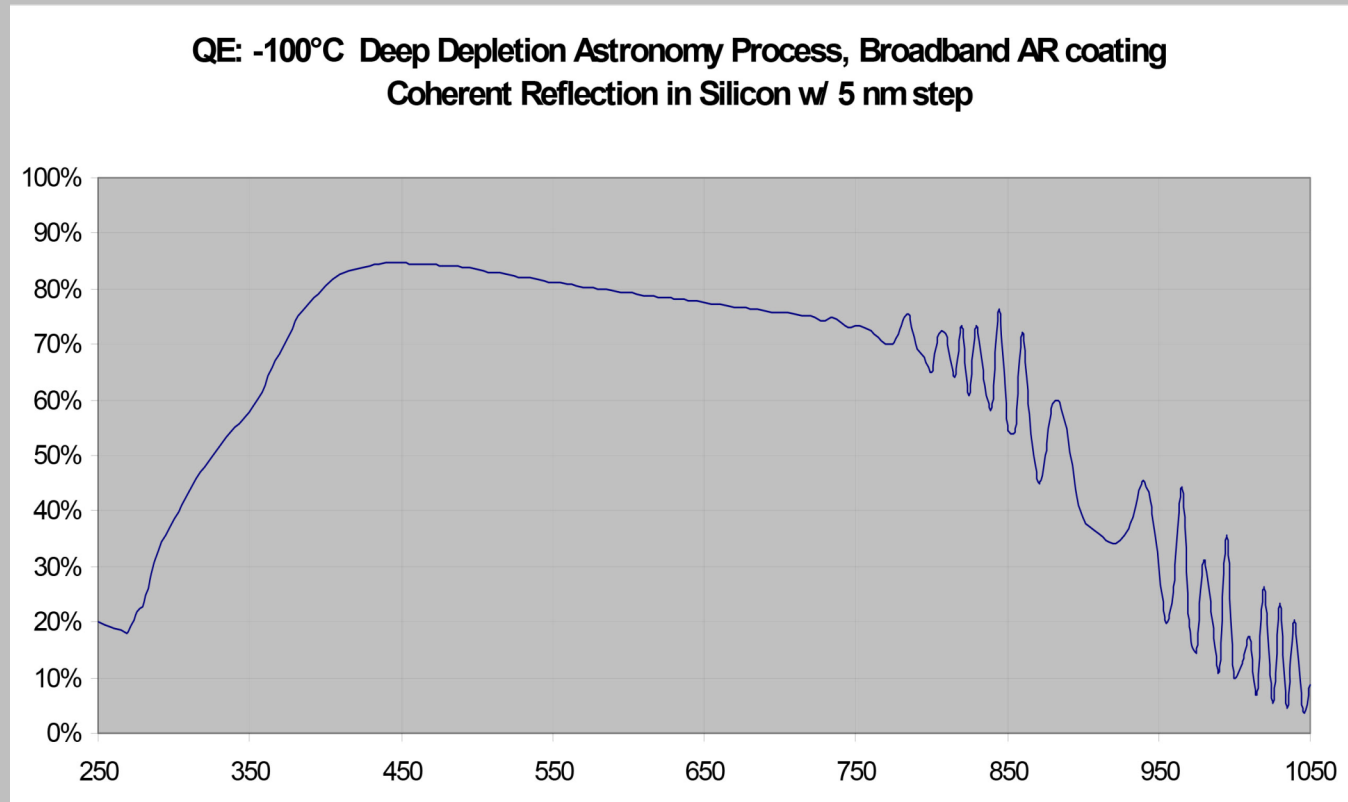
Oct 13, 2011 | 02:21 PM
Gain: 2.6377 e-/ADU
Std Dev: 1.9465 ADU
Noise: 5.1342 e-

Quantum Efficiency



- Range of coatings shown above
- Expected to go with the Astro Broadband (blue curve)

Fringing



- Plot of deep depletion fringing for CCD231
- Optional fringe suppression option to decrease fringing to 2%